# 超広視野近赤外スペースミッションの検討

Feasibility Study of Very Wide-Field Near-IR Space Mission

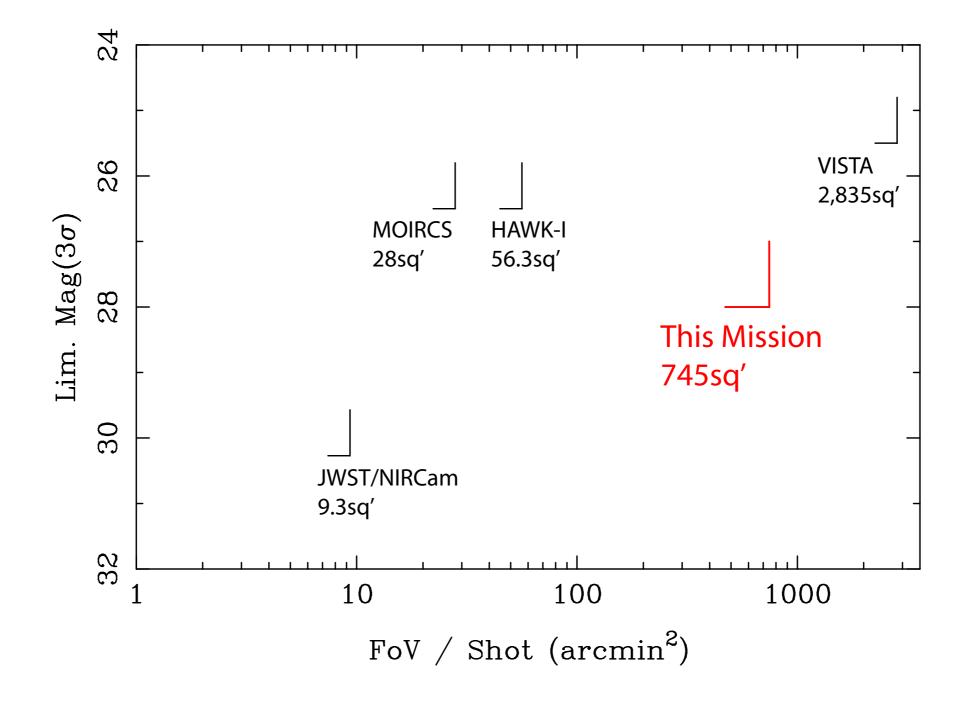
岩田 生<sup>1</sup>, 常田 佐久<sup>1</sup>, 山田 亨<sup>2</sup>, 児玉忠恭<sup>1</sup>, 近藤 善信<sup>1,3</sup> ほか (1: 国立天文台 2:東北大 3:法政大)

ポスターA24b

# Very Wide-Field Near-IR Space Mission

- 1.5m Telescope
- FoV: ~30'φ(700 arcmin<sup>2</sup>/shot)
- Wavelengths: 1-2.5μm (possibly <5μm)
- Limiting Magnitude: K~28AB
  - Unprecedented Depth (Unreachable from ground)
- Dedicated Mission: Superb Wide Survey Area (~100 sq. degrees)

## Very Wide-Field Near-IR Space Mission



# Scientific Objectives

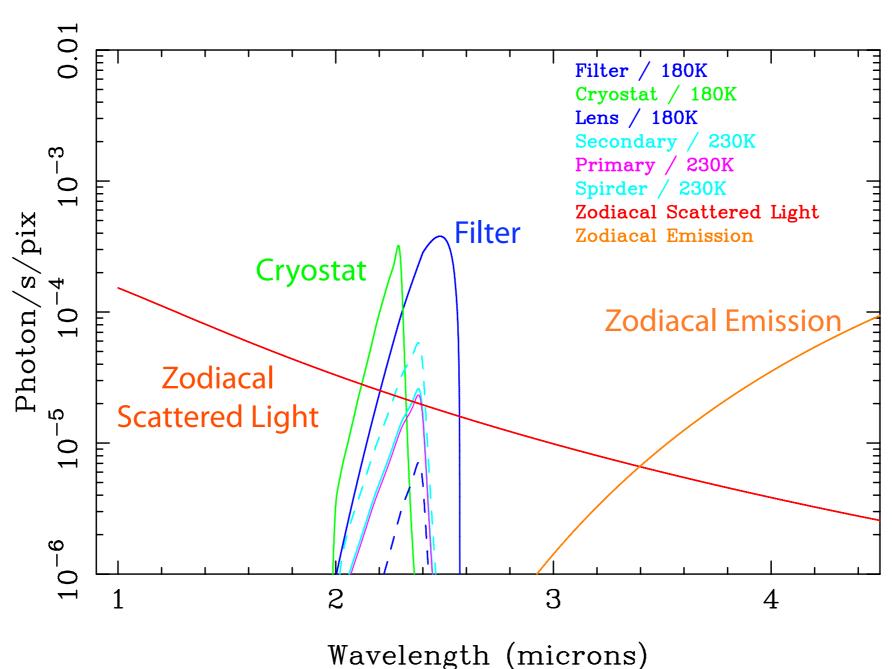
- Direct Extensions from Results of the Subaru Telescope
- Complimentary with Other Future Projects
- Detection of "First Galaxies" (z=7-14)
- Census of the Reionization Epoch
- Unbiased View of Galaxy Assembly at z=1-3
  - Stellar Mass based Large Scale Structures
  - Stellar Mass Function at Very Low-end

## Feasibility

## **Thermal Emission**

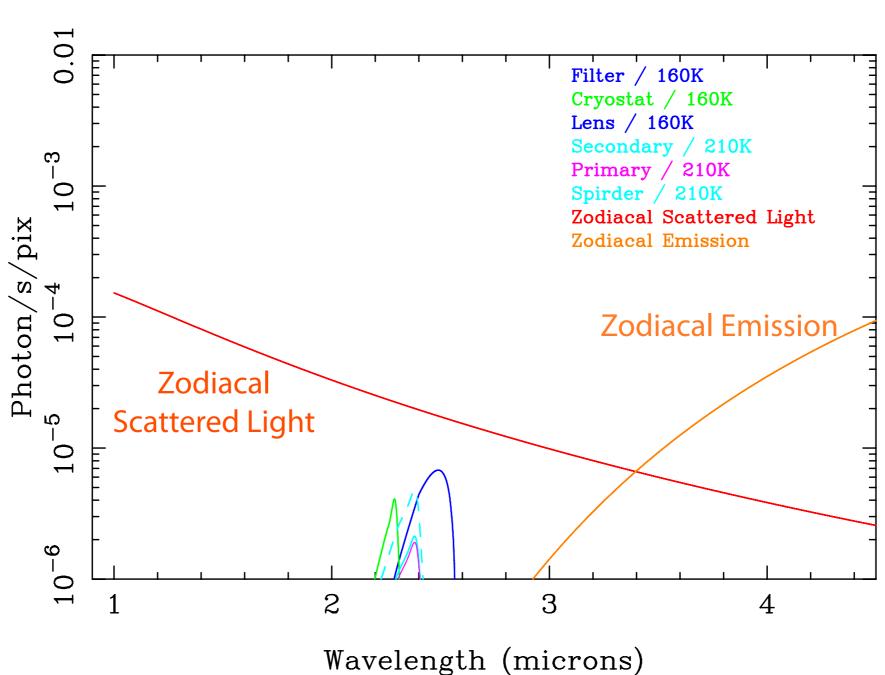
- Thermal Emission from Instruments
  - Primary and Secondary Mirrors, including Spider
  - Filter, Three Corrector Lenses
  - Cryostat

#### K-band Thermal Emission, Cryostat=180K, Mirror=230K



TEST03: 1.5m K CRY0=180 MIRROR=230

#### K-band Thermal Emission, Cryostat=160K, Mirror=210K



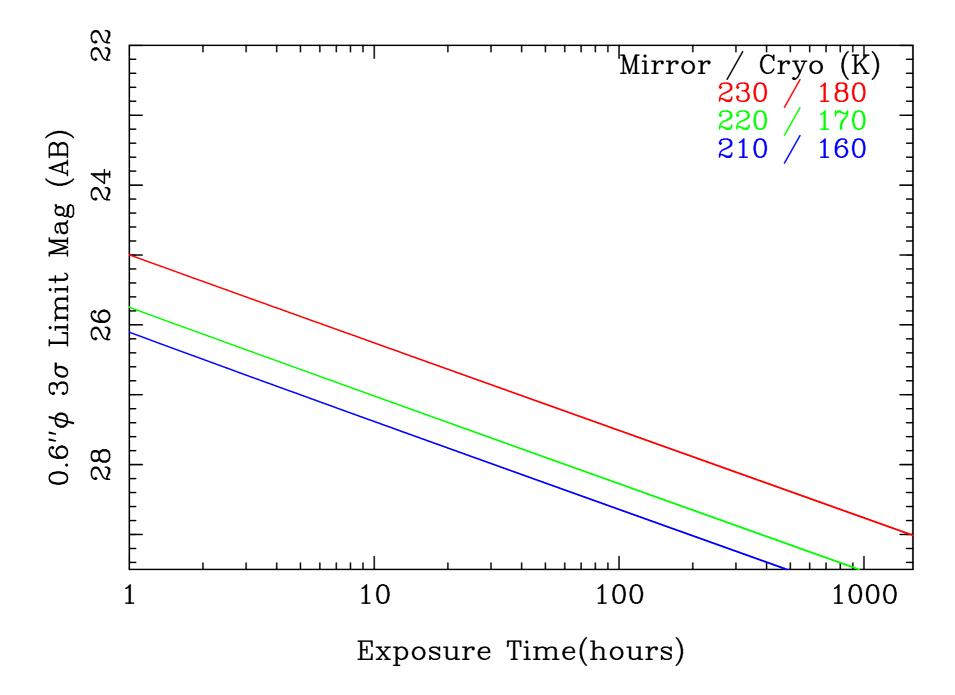
TEST02: 1.5m K CRY0=160 MIRROR=210

## Expected Sensitivity

- Backgrounds
  - Zodiacal Lights (Scatter, Thermal)
  - Thermal Emission from the Telescope and Instruments
- Dark: 0.05e-/sec/pix
- Read-out Noise: 10e-
- System Efficiency: ~30%
- Object: *fv*=flat, FWHM=0.5", Aperture=0.6" (at  $\lambda$ <2.5 $\mu$ m)

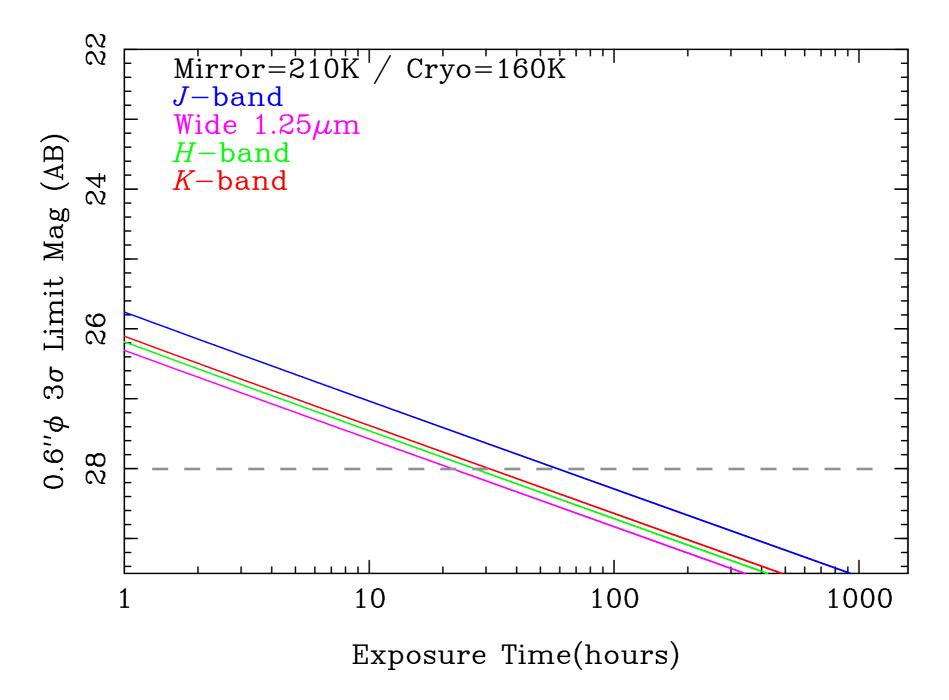
### Expected Sensitivity at K-band

K-band Limiting Mags for 1.5m Space Telescope



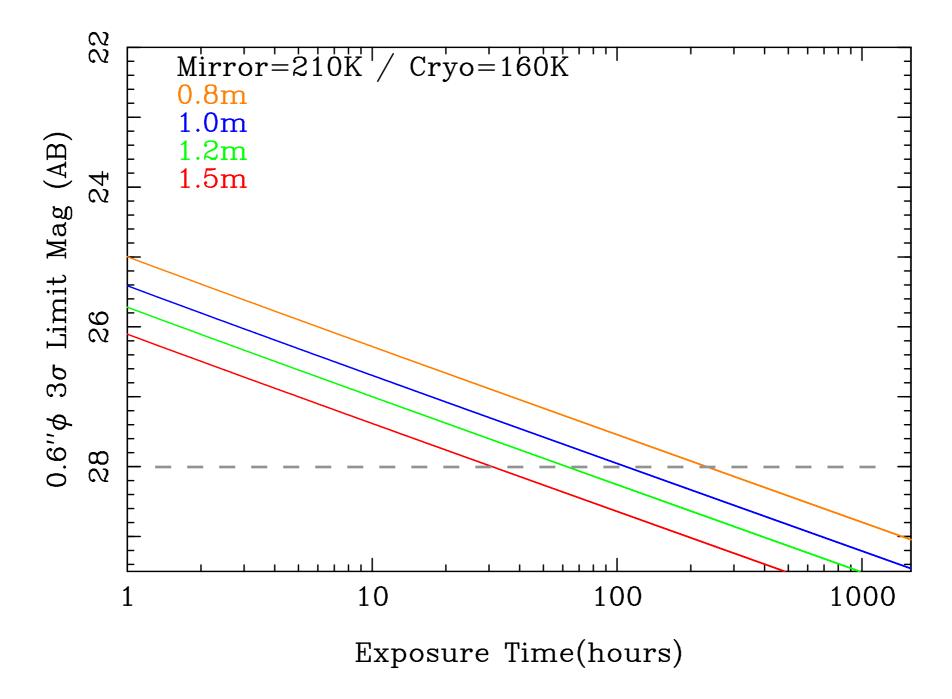
### **Expected Sensitivity with Different Filters**

Limiting Mags for 1.5m Space Telescope



#### **Expected Sensitivity with Different Mirror Sizes**

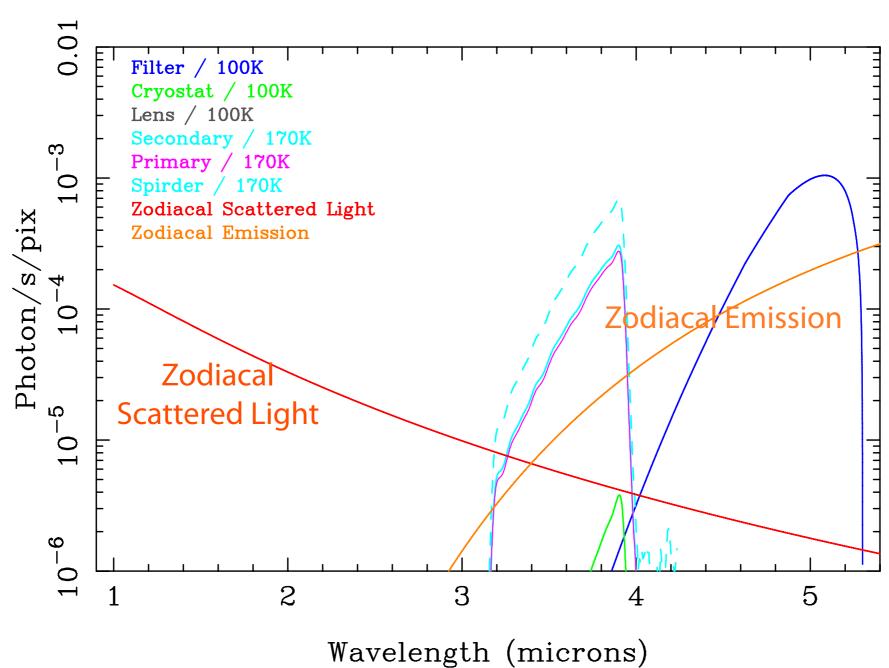
K-band Limiting Mags for Cooled Space Telescope



# Summary

- With 1.5m Telescope and Instruments at 160K, We will be able to obtain Unprecedented Deep Near-IR Images (~28AB mags) for ~100 sq. degrees.
- Thermal Design will be the Key.
- R&D including Detailed Optical Design are in Progress.
- see Poster A24b for details.

#### 3.6µm Thermal Emission, Cryostat=100K, Mirror=170K



TEST06: 1.5m 3.6um CRY0=100 MIRR0R=170

#### Expected Sensitivity at 3.6-4.5µm

 $3.6\mu m/4.5\mu m$  Limiting Mags for 1.5m Space Telescope

