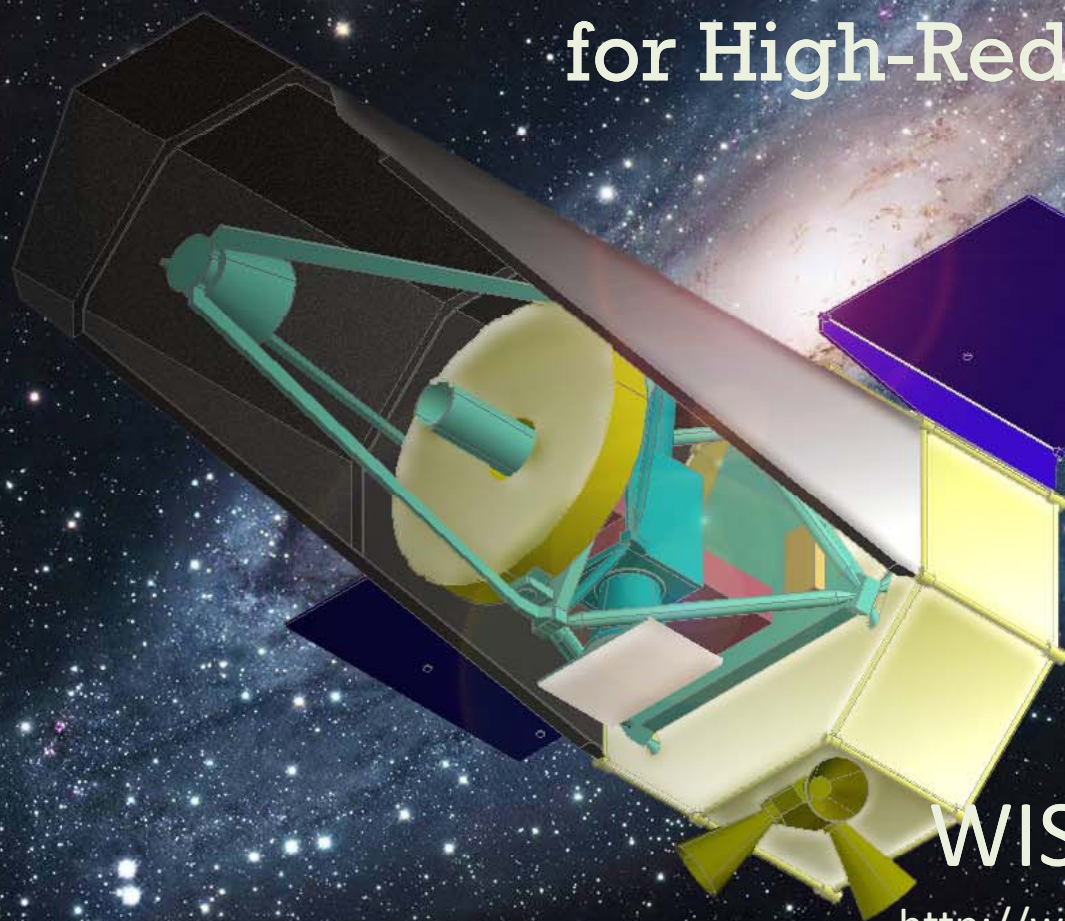


WISH

Wide-field Imaging Surveyor
for High-Redshift



1. WISH: Brief Summary
2. WISH: Science Goals
3. WISH: Survey Strategy
4. WISH: Development

WISH Working Group

<http://www.wishmission.org/en/index.html>

M31 Phot: R.Gendler

WISH WG Members

Toru Yamada, Chihiro Tokoku (Tohoku University)

Ikuru Iwata, S.Tsuneta,

T.Morokuma, T.Kodama, Y.Komiyama (NAOJ)

H.Matsuhara, T.Wada, Y.Oyabu (JAXA/ISAS)

H.Sugita, Y.Sato (JAXA)

K.Ohta, K.Yabe , Tsutsui (Kyoto University)

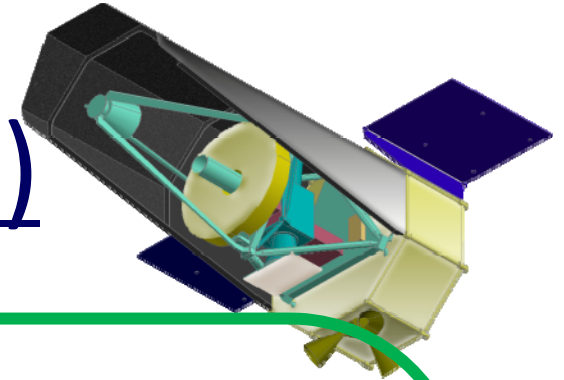
M.Doi, N.Yasuda (University of Tokyo)

N.Kawai (TiTEC), Yonetoku (Kanazawa U),

A.Inoue (Osaka Sangyo U.)

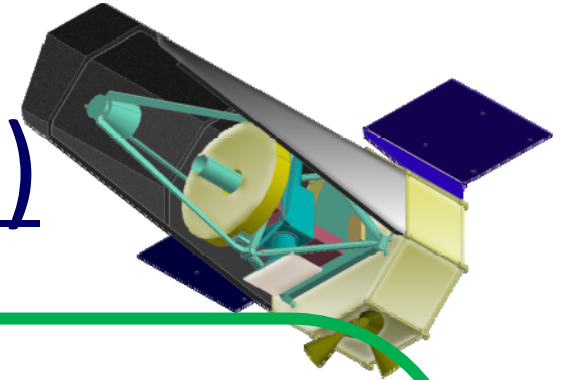
Y.Ikeda (Photocoding), S.Iwamura (M.R.J)

WISH Brief Summary (1)



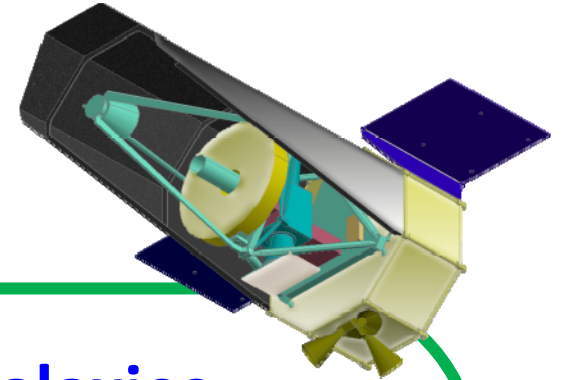
- NIR Deep and Wide-field Imaging Surveyor
- Exploring the 1st generation galaxies
- Dedicated, $\sim 100 \text{ deg}^2$, 28AB ($\sim 25 \text{ nJy}$)
- Concept developed under JAXA/ISAS WG
(the WG was selected in Sept 2008)
to be launched in late 2010's (NET2017)

WISH Brief Summary (2)



- 1-5 μm wavelength range
- 1.5m diameter telescope
- Very Wide-Field Imager
 - ~1000 arcmin^2 FoV
- pixel scale: 0.15" / 18 μm (f/16)
- Cooled to < 100K (telescope)
- SE-L2, JAXA HIIA

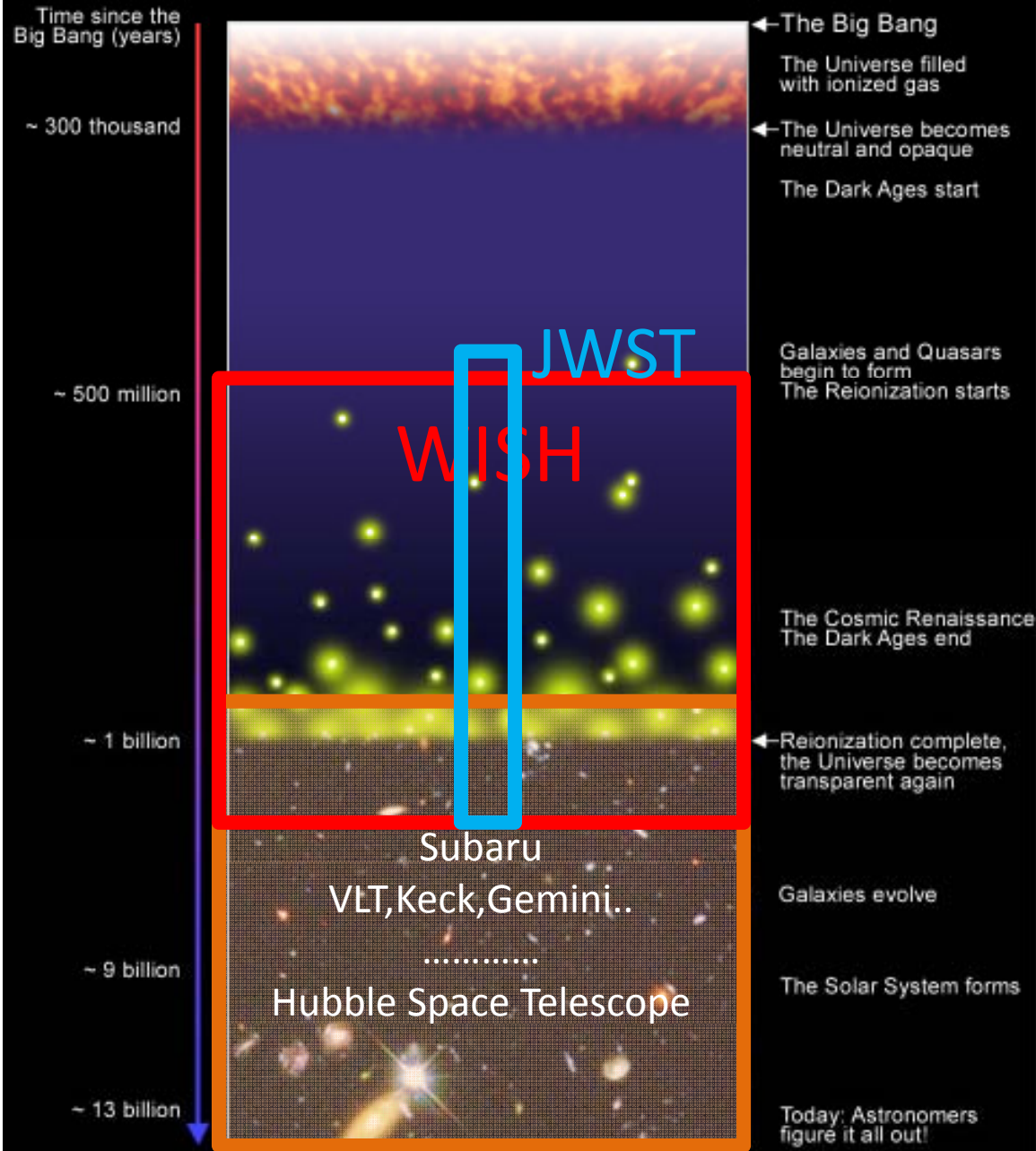
WISH Science Goals (1)



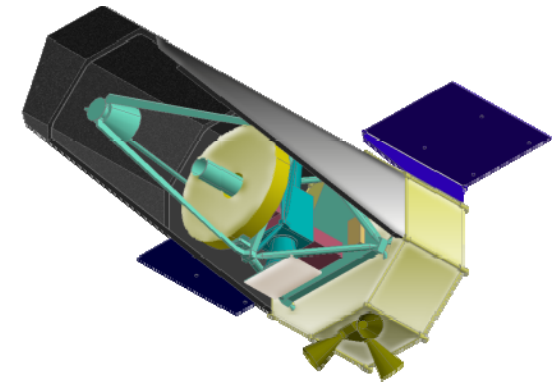
- **Exploring the Ultimate Frontier of Galaxies**
Detection of 1st-generation galaxies and studying cosmic reionization over $z=7-15$
- **NIR search and light curves for type-Ia SNe**
History of cosmic expansion and Dark Energy
- Transients: high- z GRB, luminous SNe
- Huge statistics and New discovery

What is the Reionization Era?

A Schematic Outline of the Cosmic History



S.G. Djorgovski et al. & Digital Media Center, Caltech

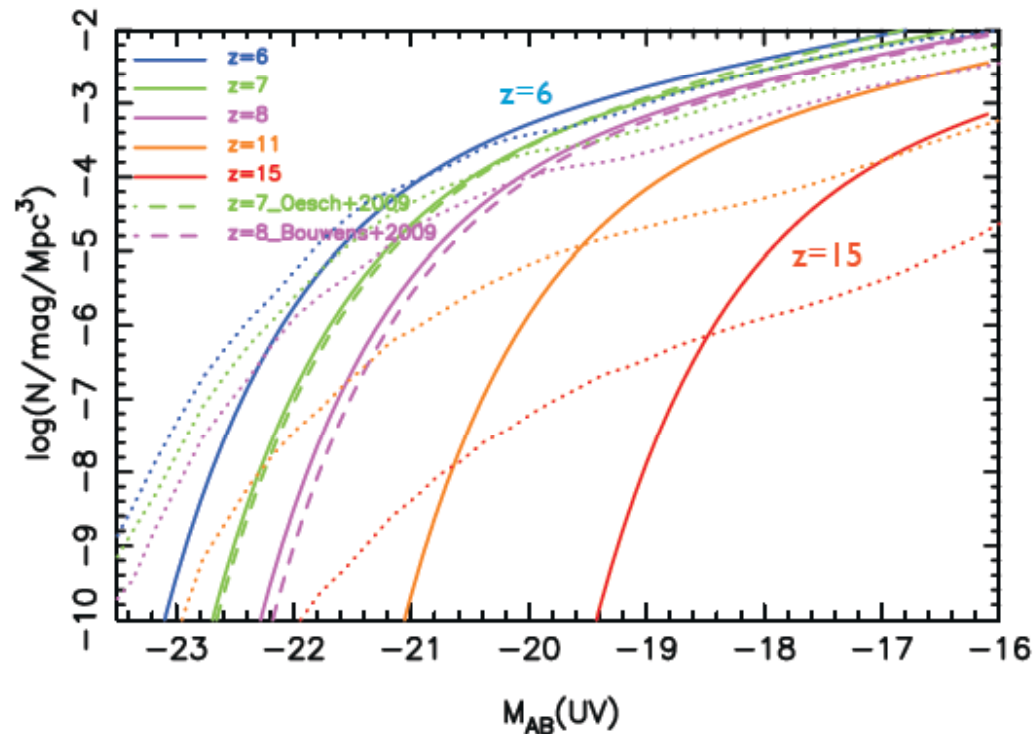


Redshift
 $z \sim 15$

Redshift
 $z \sim 6$

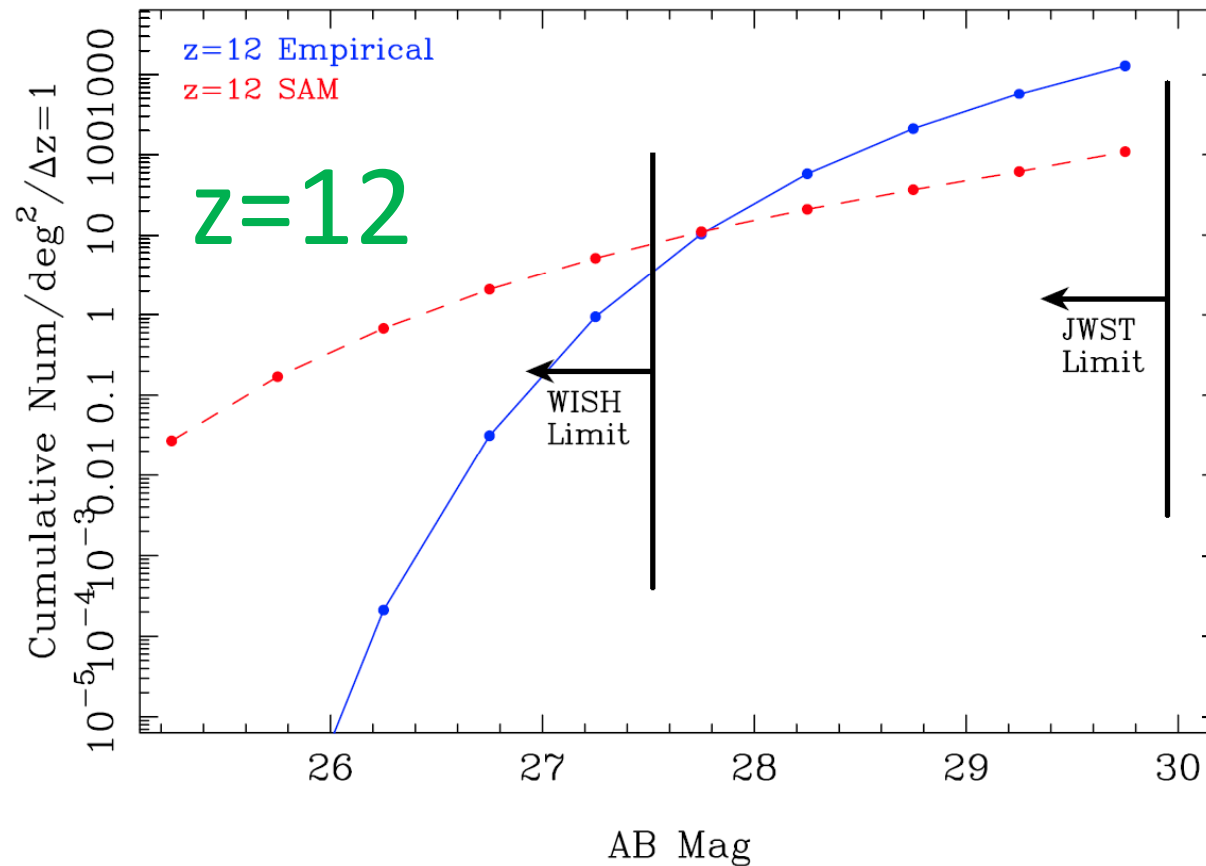
How deep, how large area should we observe?

Observed (z=6-8, HST WFC3 **dashed lines**) and
Predicted (z=6-15) UV Luminosity Function of Galaxies



Solid lines: empirical expectation (extrapolation) from z=6-8 luminosity function
Dotted lines: expectation based on galaxy models (semi-analytic treatment)

In the observers' frame...



Limit of
AB>27-28 is
needed

—
Empirical
luminosity evolution

- - -
Semi-analytic
Model prediction

FoV JWST NIRCam 2.2'x2.2' x 2ch (per filter)
~ 2.8x10⁻³ deg²

WISH Science Goals

Expected number of the *observed*
very high-redshift galaxies

		Number Density [objects per 1 deg ²] for AB < 28.0			
	redshift	No Evolution	Empirical	SAM	DMH
1.0μm-drop	8-9	4,000	1,700	630	850
1.4μm-drop	11-12	2,400	100	50	4.1
1.8μm-drop	14-17	1,200	0.72	1.1	0.003

Numbers for 1 deg² , <28AB

Galaxies bright enough for deep spectroscopy
with ELT + AO spectrograph

What is your WISH? Requirements

Limiting magnitude

(at least) **AB~28 mag** within a reasonable amount of time

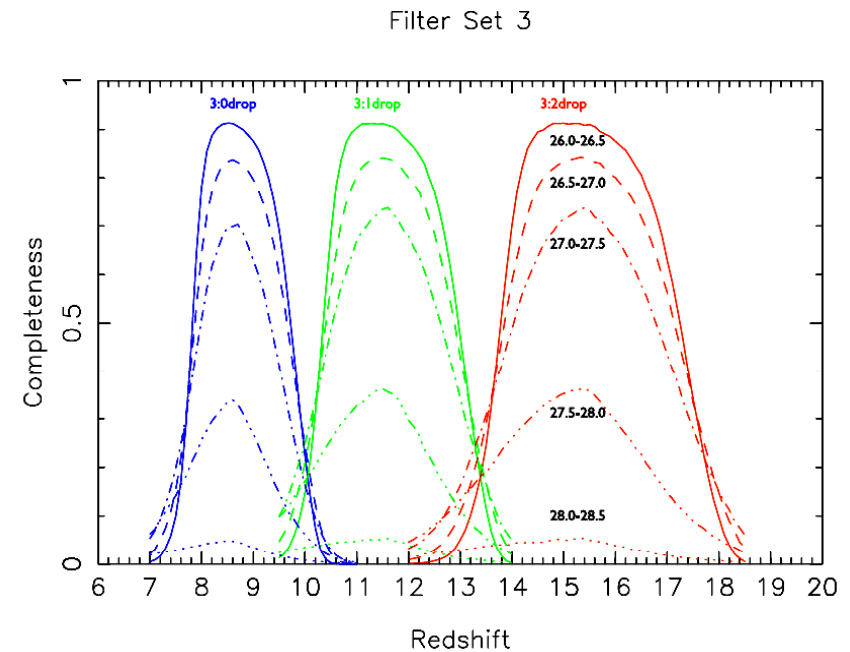
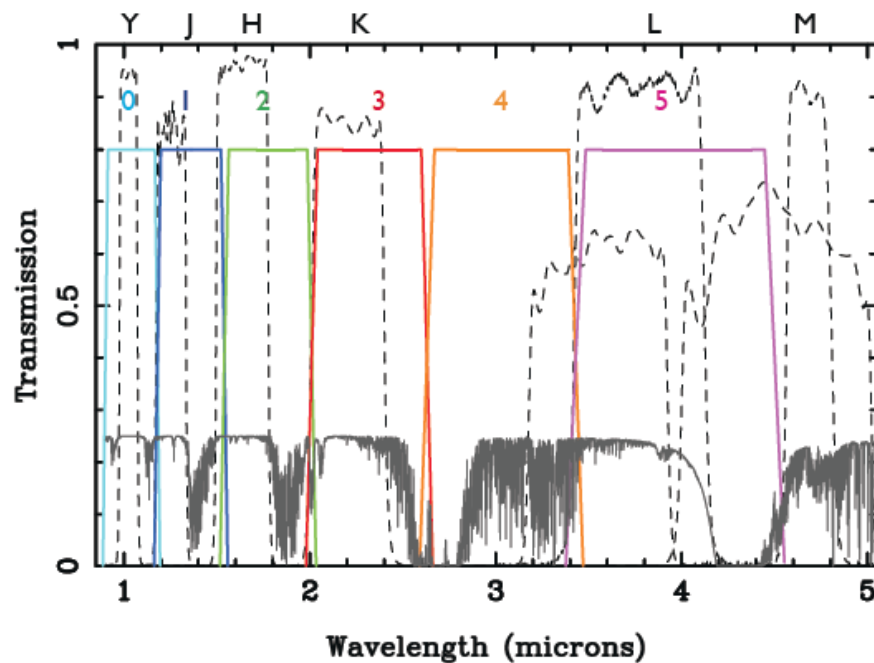
- telescope aperture **1.5m**
- image size / pixel scale **0.15"** / 18 μ m pix
vs. diff. lim. 0.22" at $\lambda = 1.5\mu\text{m}$
- telescope temperature **~100K**
to achieve ZL-limited observations

Field of View **~1000 arcmin²** (0.28 deg²)

survey speed **~2x** of JWST NIRCам for extended sources
such as $\Phi \sim 0.2''$ (**resolved for JWST, not resolved for WISH**)

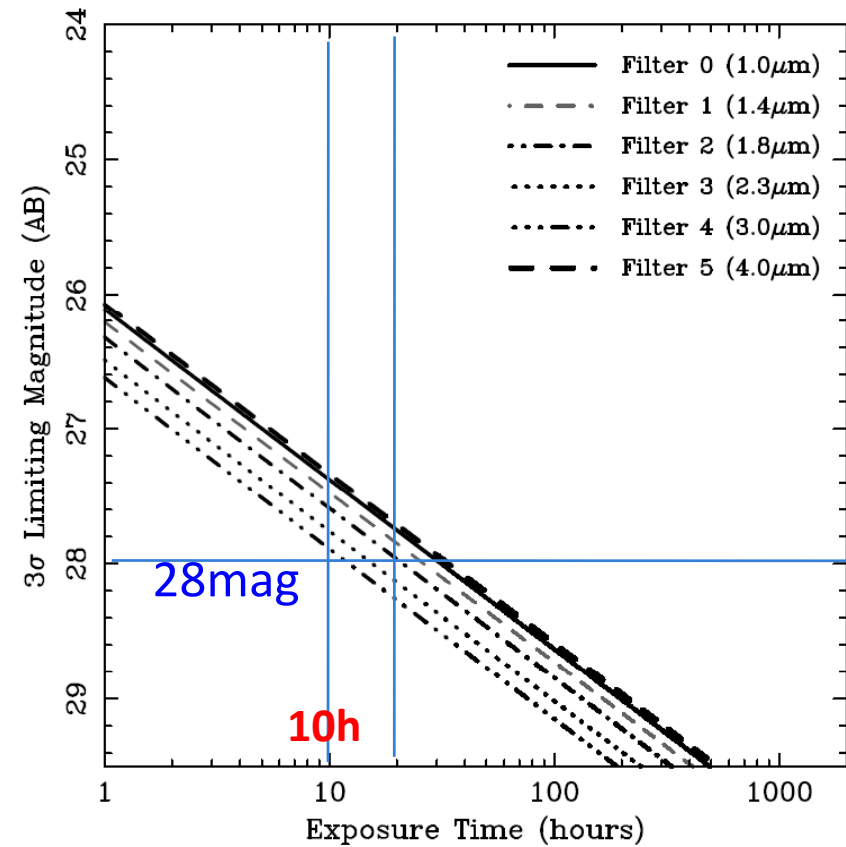
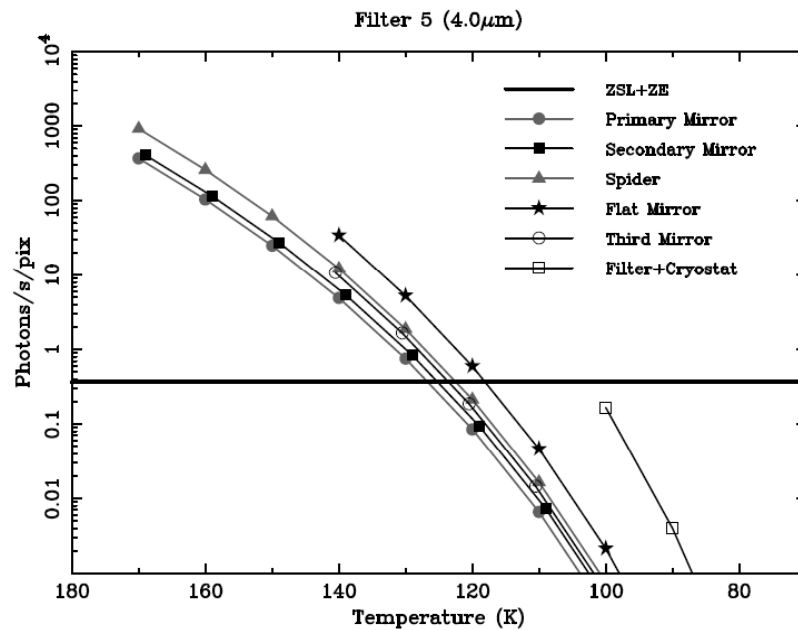
WISH: Survey Strategy

set of broad-band filters to select high- z galaxies



WISH: Survey Strategy

Telescope should be cooled to $\sim 100\text{K}$
(detector to $\sim 40\text{K}$ for $5\text{ }\mu\text{m}$)



WISH: Survey Strategy

Survey categories

	Depth (3σ) (AB mag)	Area	Example of the Filters (a plan, to be determined)
Ultra Deep Survey (UDS)	28	100 deg ²	1.4,1.8, 2.3, 3.0 μm
Multi-Band Survey (MDS)	28	10 deg ²	1.0,4.0
Ultra Wide Survey (UWS)	24-25	1000 deg ²	1.4, 1.8, 2.3
Extreme Survey	29-30	0.25 deg ²	1.0, 1.4, 1.8

WISH: Survey Strategy

Surveys achieved within ~ 1000 days (50% overhead)

WISH can detect

$\sim 10^4$ galaxies at $z=8-9$,

$\sim 10^{3-4}$ galaxies at $z=11-12$,

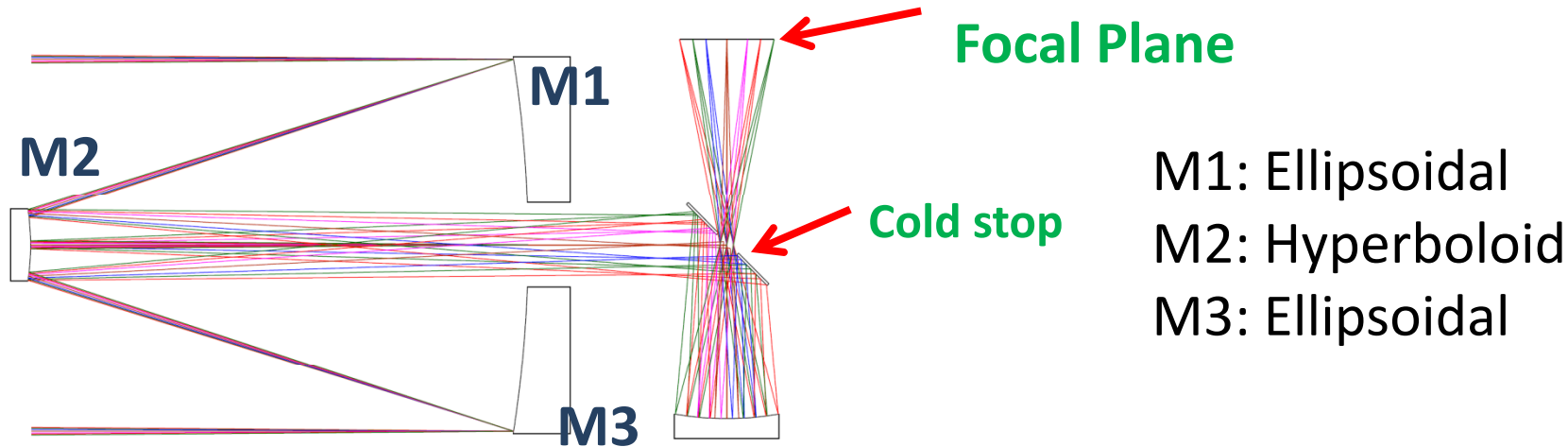
and

$\sim 50-100$ galaxies at $z=14-17$

Many of them are feasible spectroscopic targets

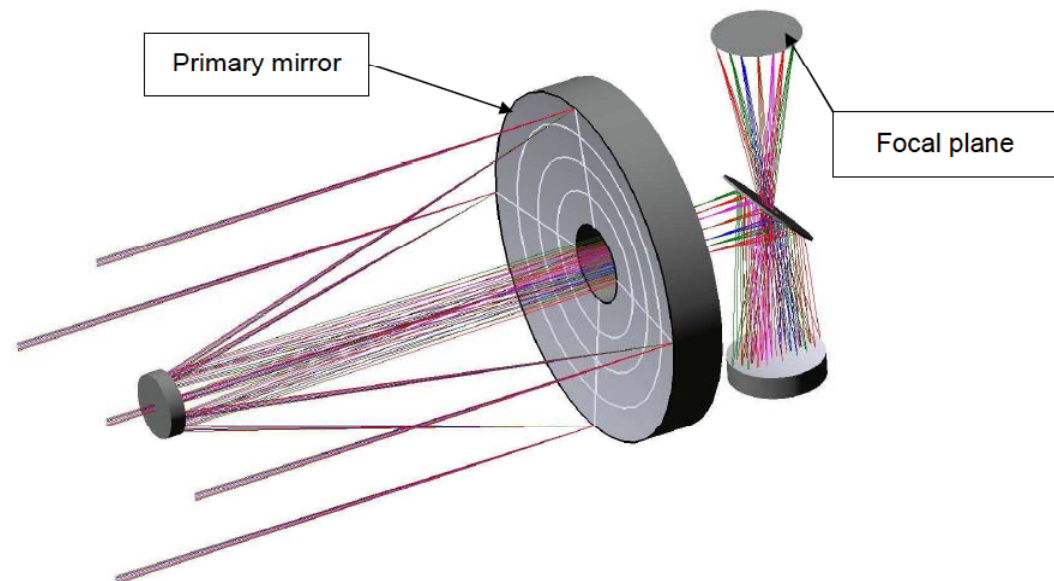
WISH Development

WISH Development: Optical Layout

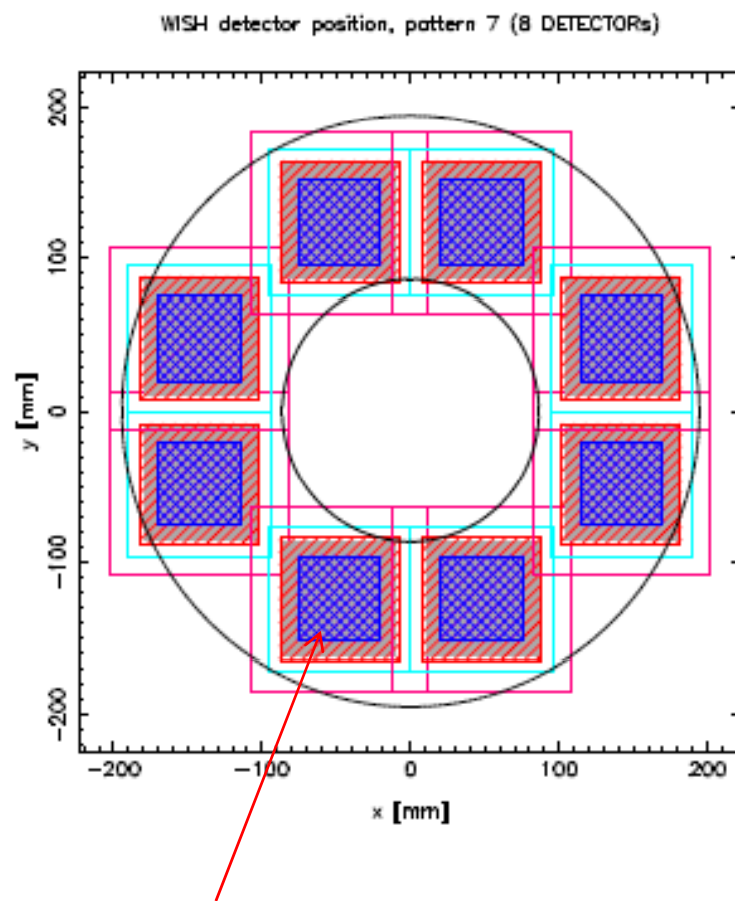


- Very flat focal plane
- Diffraction-limited images to $\phi \sim 50'$ at $1-5 \mu\text{m}$

Yuji Ikeda et al.
(photocoding)



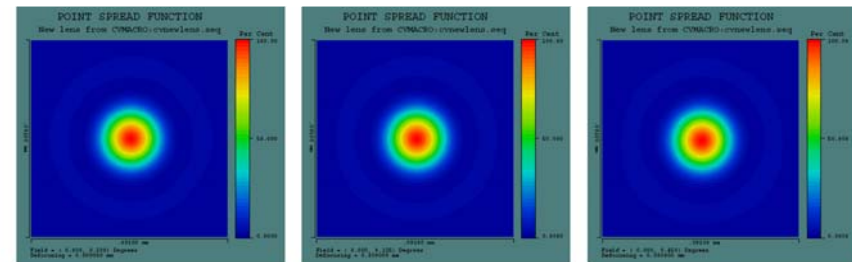
WISH Development: Optical Layout



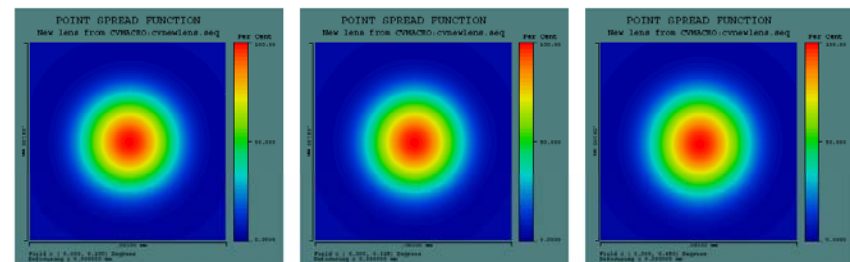
4kx4k (4 x 2k x 2k) array

Simulated Point Spread Function
at radius 0.2, 0.325, and 0.45 deg

at $\lambda \sim 1.3 \mu\text{m}$ $\text{SR} > 0.99$

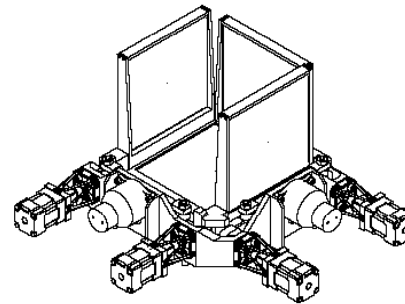
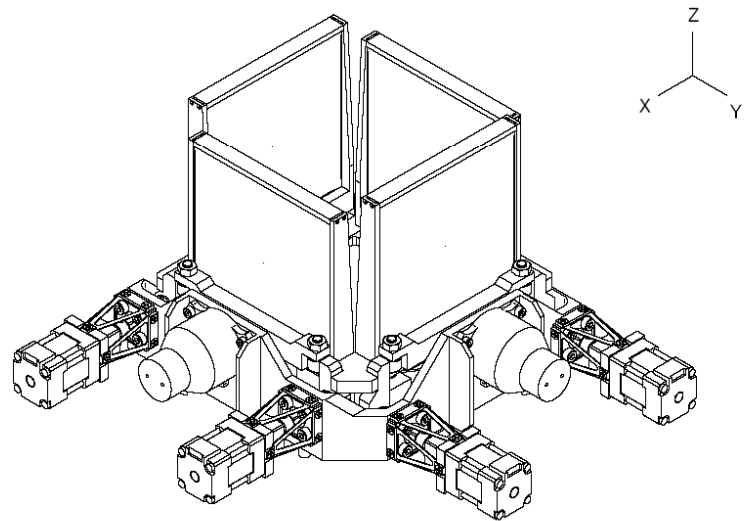


at $\lambda \sim 2.2 \mu\text{m}$ $\text{SR} > 0.99$

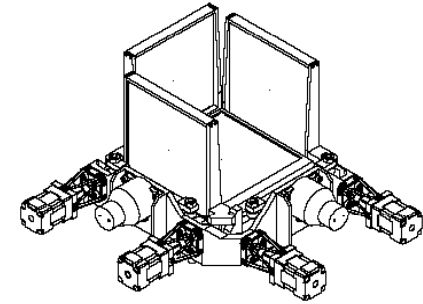


WISH Development: Filter Exchanger

Flip-type filter exchanger



CASE A

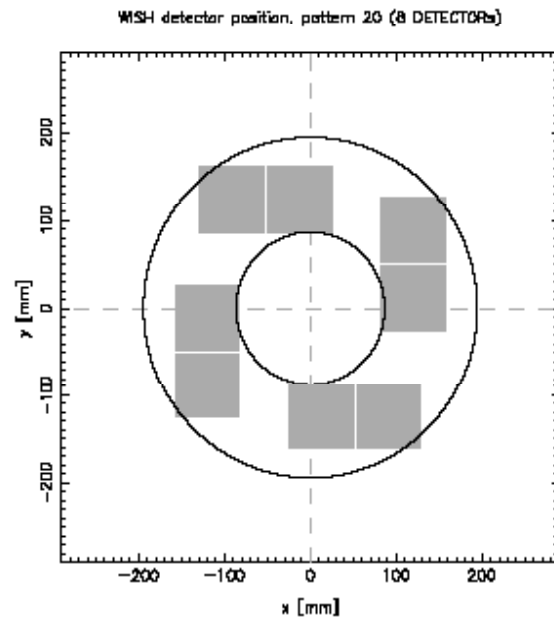


CASE B

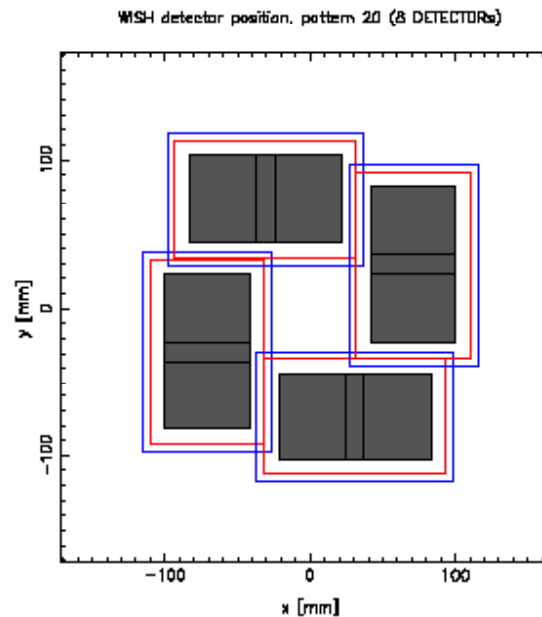
in use

WISH Development: Filter Exchanger

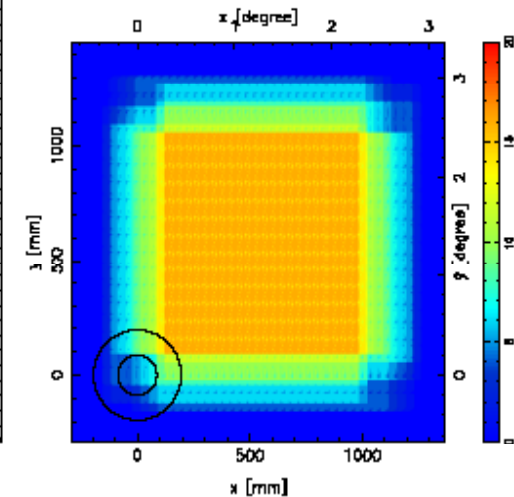
Detector Focal Plane



Filter location



Dithering Pattern

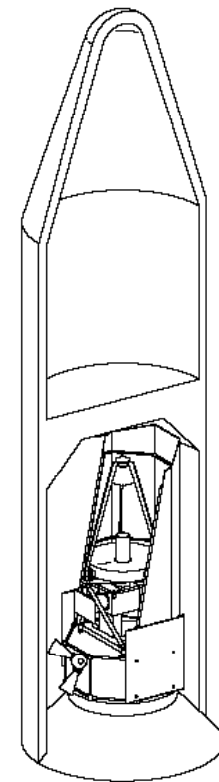
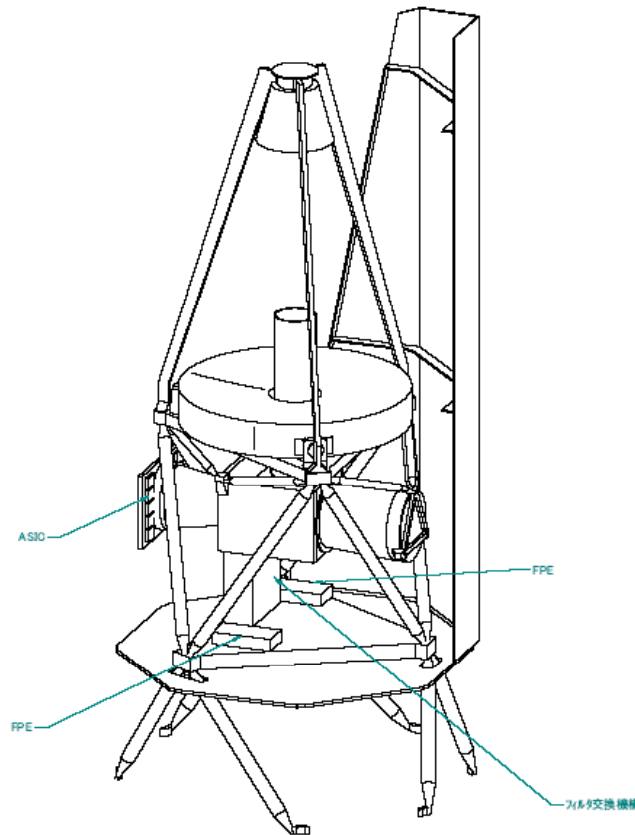
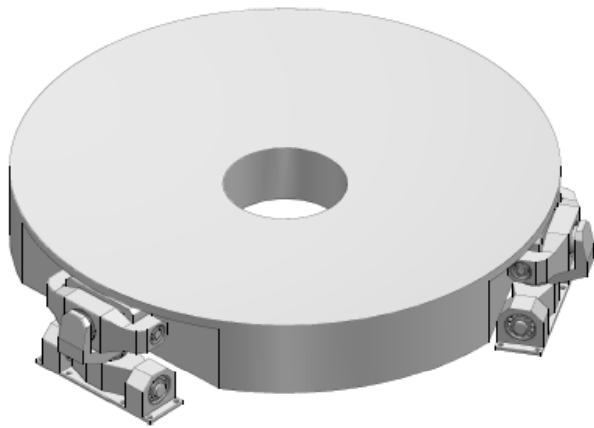


- 1 filter exchanging unit for a two (4k x 4k) arrays
- 8 filters for 1 unit

WISH Development: Telescope Structure

1.5m Primary Mirror

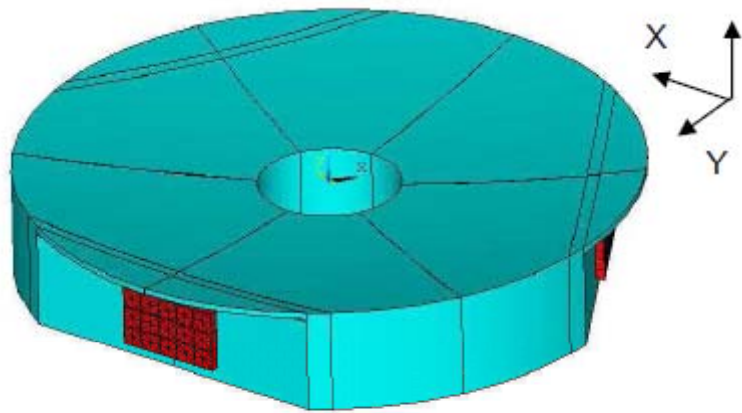
- ultra low-expansion glass material
- light weighted (< 200kg)



H-IIA (4/4D-LC)

Supported by CFRP structure

WISH Development: Telescope Structure

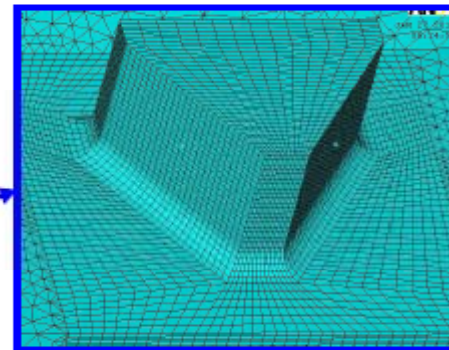
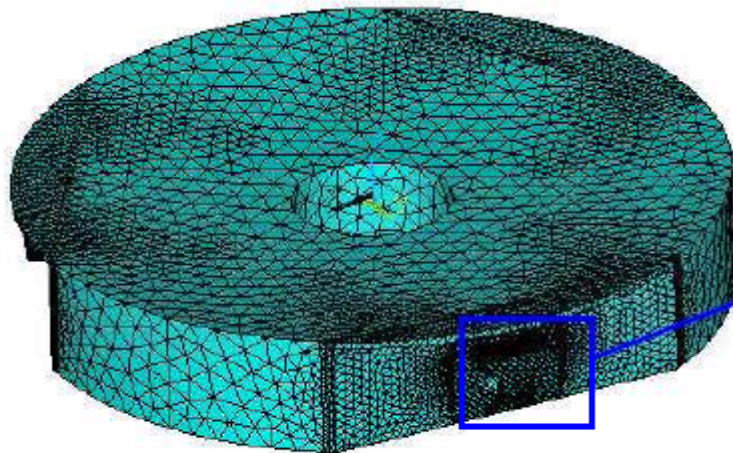


Global FEM

Mirror fixation

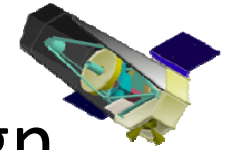
- launch load
- thermal road (to 100K)

Bonding / clumping
solutions studied

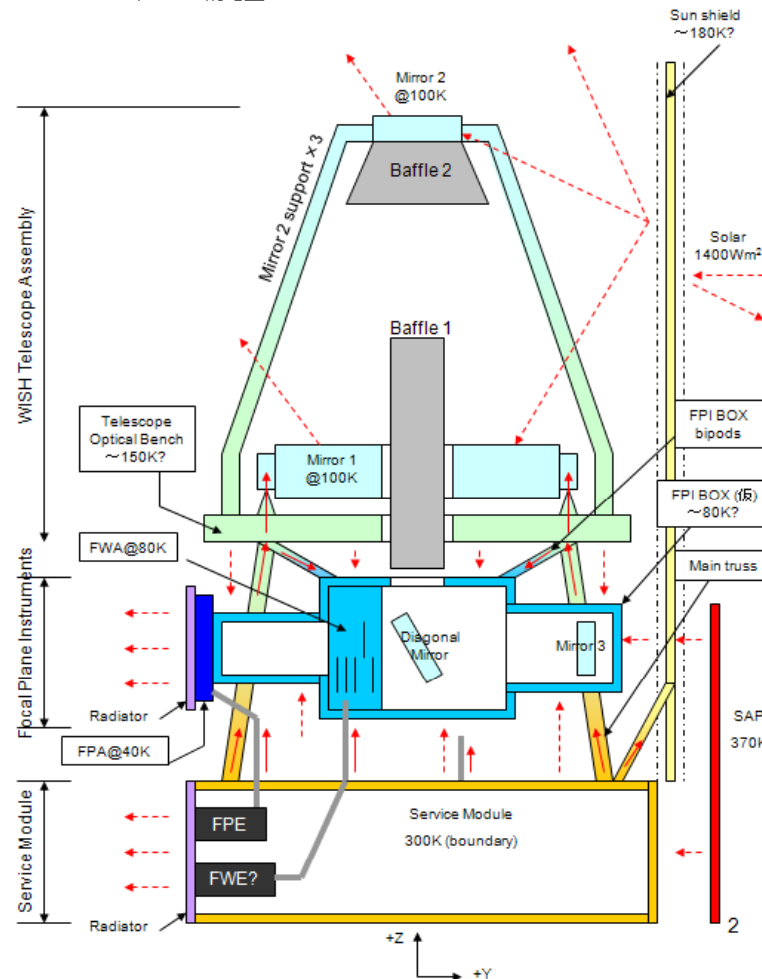


w/ SAGEM

WISH Development: Preliminary Thermal Design



■ WISH ヒートフロー概略図



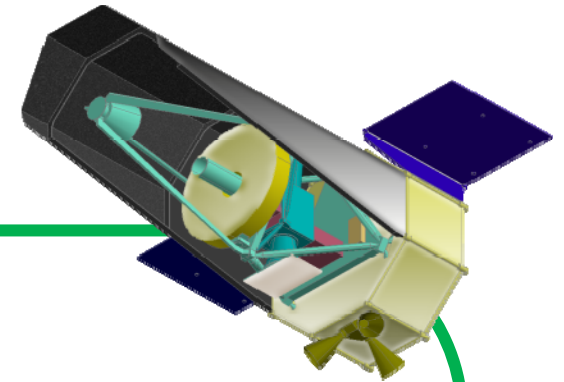
Passive cooling

- Telescope 100K
- FPI Box 80K
- Detector 40K (radiator)

Preliminary Thermal Design
On-going

	achieved	requirement
M1	79K	~100K
M2	132K	~100K
FPI BOX	95K	~80K

Summary



WISH

- NIR **Deep** and **Wide-field** Imaging Surveyor
- **1.5m** aperture, **0.15''**/pix
- Exploring the 1st generation galaxies
- Dedicated, **$\sim 100 \text{ deg}^2$, 28AB ($\sim 25 \text{ nJy}$)**
- **$\sim 10^4$ galaxies at $z=8-9$, $\sim 3-6 \times 10^3$ at $z=11-12$,
and $\sim 50-100$ galaxies at $z=14-17$**
- Concept developed under JAXA/ISAS WG
to be launched in late 2010's