# SNIa Cosmology with WISH satellite Nao Suzuki (Lawrence Berkeley Lab)

§1: Introduction to Type Ia Supernova
§2: SNIa Cosmology Today
§3: SNIa Challenges (systematic errors) New Standard Stars for IR
§4: SNIa Cosmology with WISH

### Supernova Type Ia : Standard Candle Phillips Relation (1993)









Physics Today (2003)

### Discovery of Dark Energy (SCP 1997)



Gerson Goldhaber reported non-zero  $\Lambda$ on 24th in Sep '97





Scale of the Universe

### Accelerating Universe

#### Expansion History of the Universe



### HST Cluster SN Survey

### Nao Suzuki & SCP PI: Saul Perlmutter







### SNIa Cosmology Today (2011)

Supernova Cosmology Project (Suzuki et al. arxiv:1105.3470)



## $\Lambda$ Today

Combination of SNe with: BAO (Percival et. al., 2010) CMB (WMAP data, 2011) For a flat Universe:

LCDM:  $\Omega_m = 0.271 \pm 0.012(\text{stat}) \pm 0.014(\text{sys})$ with curvature: oLCDM:  $\Omega_k = 0.002 \pm 0.005(\text{stat}) \pm 0.005(\text{sys})$ 



### w=P/ $\rho$ : equation of state Q. Is *w* -1?

wCDM:  $w = -1.008 \pm 0.052(stat)$   $-1.013 \pm 0.070(sys)$ SNe + BAO + CMB ... and allowing for curvature: owCDM  $w = -1.006 \pm 0.058(stat)$   $-1.003 \pm 0.093(sys)$ with systematics

•w=-1 : cosmological constant •w=0 : matter •w=1/3: radiation  $E \propto a^{-3(1+w)}$ 



### Future of SNIa Cosmology Q: Can we Reduce Systematic Errors?

 Q1 : Observational Systematic Error => SDSS stars and Perfect Black Body stars

• Q2 : SNIa is not well understood? => Spectral Library & Simulation

### The Origins of Systematic Errors Today

- I : Zero Point is not accurate enough (will be calibrated them with 300,000 SDSS F-stars)
- II : Standard Star Calibration (HST CALSPEC) is not accurate enough (will be replaced by DC WDs)



Discovery of Perfect Black Body Spectra: found as a Quasar Target Only 20 stars out of a million stars

• DR8 : 605,772 stars => 5 stars

• DR9 : 110,929 stars=> 9 stars

• DR10: 81,892 stars=> 10 stars (July 2013)







# Looking at the Future with WISH/JWST today 0.3%, but we can do 0.1%



### SNIa models 2D Simulation => Observation Kasen, Ropke, Woosely (2009)





- 2D model
- 56Ni,
- Oxygen Mass,
- Carbon Mass
- Kinetic Energy
- Reproduces LTCV









# **SNIa Rates**

- z > 1.0 is now well understood
- (Gauer et al 2011) assume rate is constant



### WISH SN Survey (100 deg) 2000 SNe (3-band, 30-day x 5)









### $w=P/\rho$ : equation of state Q. Is *w* constant over time? wa=0? A. No clear sign of non-zero wa, yet

Dark Energy Task Force (DETF) Figure of Merit : FoM  $\sigma(w_0)\sigma(w_a)$  $w(a) = w_0 + (1 - a)w_a$ FoM=39.3(stat)22.6(sys)Suzuki et al 2012 **SCP UNION2.1** 



# Dark Energy in 2015-2025: Conclusion & Summary

- Reducing Systematic is mandatory
- Calibration is the key
- SNIa + CMB (Planck) + BAO woul powerful combination in the next 10 years



- DES can find z > 1 SNe but needs z, y-band
- With WISH SNIa can lead Dark Energy Studies
- New Discoveries are yet to come (PopIII etc)
- WISH satellite can enhance FoM => 300

# Backup Slides

#### SN Factory, Palomar Transient Factory (PTF), SuperNovaLegacySurvery (SNLS), Suprnova Cosmology Project (SCP)



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### From Deceleration to Acceleration

